

3

INTERACTIONS

- If we want to describe interactions, we must add non-linear terms to the Lagrangian.
- We must create renormalizable theories to avoid divergencies. It can be shown that theories that involve coupling constants of dimension $(\text{mass})^{-n}$ (mass in negative power) are not renormalizable. This limits the # of possible interactions to only a few.

→ For a scalar field:

$$[\phi] = m \leadsto \alpha \phi^3 \Rightarrow [\alpha] = m^{-1}$$

$$\lambda \phi^4 \Rightarrow [\lambda] = m^0$$

$$c \phi^5 \Rightarrow [c] = m^{-1} \text{ (2)}$$

↳ We only have 2 terms possible.

3.1 Interaction representation

① Schrodinger and Heisenberg representation

- In QM, in Schrodinger representation, the evolution of the system is described by the Schrodinger equation:

$$i\partial_t \psi(t) = \hat{H} \psi(t)$$

which has the formal solution $\psi(t) = e^{-i\hat{H}t} \psi(0)$

- The observables are the averages of operators:

$$\langle \hat{B} \rangle = \langle \psi(t) | \hat{B} | \psi(t) \rangle$$

- Using the solution (2), we introduce the Heisenberg representation:

$$\langle \hat{B} \rangle = \langle \psi(0) | e^{i\hat{H}t} \hat{B} e^{-i\hat{H}t} | \psi(0) \rangle$$

$$\equiv \langle \psi(0) | B_H(t) | \psi(0) \rangle$$

⊙ Construction of the interaction representation:

→ Consider the Hamiltonian as a sum of the free part $\hat{H}_0$ and the interaction $\hat{H}_I$. The $\ddot{S}$ become:

$$i \partial_t \psi(t) = (\hat{H}_0 + \hat{H}_I) \psi(t) \quad \hat{H}_0 = \int d^3k \{ \omega_k a_k^\dagger a_k \}$$

? → We consider solution of the form $\psi(t) = e^{-i\hat{H}_0 t} \psi_I(t)$
Inserting, we get:

$$i \partial_t \psi = \hat{H}_0 e^{-i\hat{H}_0 t} \psi_I + i e^{-i\hat{H}_0 t} \partial_t \psi_I = \hat{H}_0 e^{-i\hat{H}_0 t} \psi_I + \hat{H}_I e^{-i\hat{H}_0 t} \psi_I$$

$$\Leftrightarrow i \partial_t \psi_I = e^{i\hat{H}_0 t} \hat{H}_I e^{-i\hat{H}_0 t} \psi_I$$

DEF | The Hamiltonian of interaction in the interaction representation is
$$\hat{H}_I^I = e^{i\hat{H}_0 t} \hat{H}_I e^{-i\hat{H}_0 t}$$

→ We then have $i \partial_t \psi_I(t) = \hat{H}_I^I \psi_I$ / $\ddot{S}$

→ Example: consider $\hat{H}_I = \int d^3x \phi^4(0, \vec{x})$

$$\text{Then } \hat{H}_I^I = \int d^3x e^{i\hat{H}_0 t} \phi^4(0, \vec{x}) e^{-i\hat{H}_0 t} = \int d^3x \phi^4(t, \vec{x})$$

→ Matrix elements:

$$\langle B \rangle = \langle \psi(t) | \hat{B} | \psi(t) \rangle = \langle \psi_I(t) | e^{+i\hat{H}_0 t} \hat{B} e^{-i\hat{H}_0 t} | \psi_I(t) \rangle$$

$$\langle B \rangle = \langle \psi_I | \hat{B}_I | \psi_I \rangle$$

↳ In interaction representation, operators are function of the free fields

⊙ Evolution operator:

→ We invert $\otimes$: $\psi_I(t_2) = e^{+i\hat{H}_0 t_2} \psi(t_2) = e^{+i\hat{H}_0 t_2} e^{-i\hat{H}_0 t_2} \psi(0)$

$$\Rightarrow \psi(0) = e^{+i\hat{H}_0 t_1} e^{-i\hat{H}_0 t_1} \psi_I(t_1)$$

$$\Rightarrow \psi_I(t_2) = e^{i\hat{H}_0 t_2} e^{-i\hat{H}(t_2-t_1)} e^{-i\hat{H}_0 t_2} \psi_I(t_1) \\ = U(t_1, t_2) \psi_I(t_1)$$

This expression is not convenient for practical calculations

3.2 Evolution in interaction representation

→ We want to find the explicit expression for the evolution operator in the interaction representation.

① Integral form of the Schrödinger equation:

→ We had $i\partial_t \Psi_I = \hat{H}_I(t) \Psi_I(t)$. In the form of an integral equation,

$$\Psi_I(t) = \Psi_I(t_0) - i \int_{t_0}^t dt' \hat{H}_I(t') \Psi_I(t')$$

↳ We can solve this iteratively. Let's substitute in the r.h.s.:

$$\Psi_I(t) = \Psi_I(t_0) - i \int_{t_0}^t dt' \hat{H}_I(t') \Psi_I(t_0) + (-i)^2 \int_{t_0}^{t'} dt'' \hat{H}_I(t'') \Psi_I(t'')$$

↳ One more time:

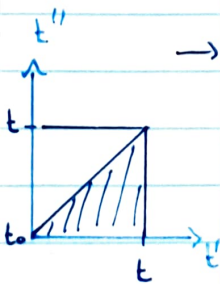
$$\begin{aligned} \Psi_I(t) = & \Psi_I(t_0) - i \int_{t_0}^t dt' \hat{H}_I(t') \Psi_I(t_0) \\ & + (-i)^2 \int_{t_0}^{t'} dt'' \hat{H}_I(t'') \Psi_I(t_0) \\ & + (-i)^3 \int_{t_0}^{t''} dt''' \hat{H}_I(t''') \Psi_I(t_0) + \dots = U(t, t_0) \Psi_I(t_0) \end{aligned}$$

$$\text{with } U(t, t_0) = 1 - i \int_{t_0}^t dt' \hat{H}_I(t') + \dots + (-i)^n \int_{t_0}^t dt' \dots \int_{t_0}^{t_{n+1}} dt_{n+1} \hat{H}_I(t_{n+1}) + \dots$$

② Complex analysis and time ordering:

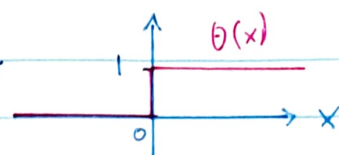
→ Consider individual terms of the series $U(t, t_0)$ (ex: 2nd term):

$$(-i)^2 \underbrace{\int_{t_0}^t dt' \hat{H}_I(t') \int_{t_0}^{t'} dt'' \hat{H}_I(t'')}_{t'' \leq t'} = (-i)^2 \int_{t_0}^t dt \int_{t_0}^t dt'' \hat{H}_I(t') \hat{H}_I(t'') \Theta(t' - t'')$$



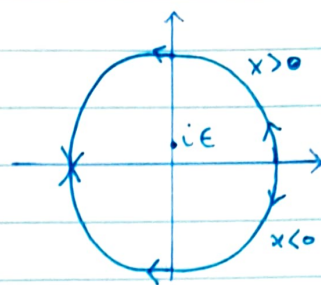
DEF The Heaviside step function θ is defined by

$$\begin{cases} \theta(x) = 0 & \text{at } x < 0 \\ \theta(x) = 1 & \text{at } x \geq 0 \end{cases}$$



Equivalently,

$$\theta(x) \equiv \frac{1}{2\pi i} \int_{-\infty}^{\infty} \frac{d\lambda}{\lambda - i\epsilon} e^{i\lambda x}$$



→ Check that $\partial_x \theta = \delta(x)$:

$$\begin{aligned} \frac{d\theta(x)}{dx} &= \frac{1}{2\pi i} \int_{-\infty}^{\infty} i\lambda \frac{d\lambda}{\lambda - i\epsilon} e^{i\lambda x} \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} d\lambda e^{i\lambda x} = \delta(x) \end{aligned}$$

DEF Let $\hat{A}(x^\mu), \hat{B}(x^\mu)$ be 2 operators. The time-order product T is defined as follow:

$$T[\hat{A}(x) \hat{B}(y)] \equiv \begin{cases} \hat{A}(x) \hat{B}(y) & \text{at } x^0 > y^0 \\ \hat{B}(y) \hat{A}(x) & \text{at } x^0 < y^0 \end{cases}$$

Equivalently,

$$T[A(x)B(y)] = \theta(x^0 - y^0) \hat{A}(x) \hat{B}(y) + \theta(y^0 - x^0) \hat{B}(y) \hat{A}(x)$$

→ The 2nd terms can be now expressed as follow:

$$(-i)^2 \int_{t_0}^t dt' dt'' \hat{H}_I(t') \hat{H}_I(t'') \cdot \theta(t' - t'')$$

$$= (-i)^2 \frac{1}{2} \int_{t_0}^{t_0} dt' dt'' \left\{ \hat{H}_I(t') \hat{H}_I(t'') \theta(t' - t'') + \hat{H}_I(t'') \hat{H}_I(t') \theta(t'' - t') \right\}$$

$$= (-i)^2 \frac{1}{2} \int_{t_0}^t dt' dt'' T[\hat{H}_I(t') \hat{H}_I(t'')]$$

$$= \frac{(-i)^2}{2} T \left[\left(\int_{t_0}^t dt' \hat{H}_I(t') \right)^2 \right]$$

→ Before going for the 3rd term, let's calculate $T[A(t_1)B(t_2)C(t_3)]$:

$$\begin{aligned} T[A(t_1)B(t_2)C(t_3)] &= ABC \theta_{12} \theta_{23} + ACB \theta_{13} \theta_{32} \\ &\quad + BAC \theta_{21} \theta_{13} + BCA \theta_{23} \theta_{31} \\ &\quad + CAB \theta_{31} \theta_{12} + CBA \theta_{32} \theta_{21} \end{aligned}$$

→ Rewriting the 3rd term with the T-product:

$$\begin{aligned}
 & (-i)^3 \int_{t_0}^t dt' \hat{H}_I(t') \int_{t_0}^{t'} dt'' \hat{H}_I(t'') \int_{t_0}^{t''} dt''' \hat{H}_I(t''') \\
 &= (-i)^3 \int_{t_0}^t dt' dt'' dt''' \cdot \hat{H}_I(t') \hat{H}_I(t'') \hat{H}_I(t''') \theta(t'-t'') \theta(t''-t''') \\
 &= \frac{(-i)^3}{3!} \int_{t_0}^t dt' dt'' dt''' T[\hat{H}_I(t') \hat{H}_I(t'') \hat{H}_I(t''')] \\
 &= \frac{(-i)^3}{3!} T\left[\left(\int_{t_0}^t dt' \hat{H}_I(t')\right)^3\right]
 \end{aligned}$$

→ Generalization:

$$\begin{aligned}
 U(t, t_0) = & 1 - i \int_{t_0}^t dt' \hat{H}_I(t') + \frac{(-i)^2}{2!} T\left[\left(\int_{t_0}^t dt' \hat{H}_I(t')\right)^2\right] \\
 & + \frac{(-i)^3}{3!} T\left[\left(\int_{t_0}^t dt' \hat{H}_I(t')\right)^3\right] + \dots + \frac{(-i)^n}{n!} T\left[\left(\int_{t_0}^t dt' \hat{H}_I(t')\right)^n\right] + \dots
 \end{aligned}$$

$$U(t, t_0) = T\left[\exp\left\{-i \int_{t_0}^t \hat{H}_I(t') dt'\right\}\right]$$

DEF

The evolution operator from $-\infty$ to $+\infty$ is called the S-matrix, defined as follow:

$$\hat{S} \equiv \hat{U}(\infty, -\infty) = T \exp\left\{i \int_{-\infty}^{\infty} dt \hat{H}_I(t)\right\}$$

↳ Notation! Can't take the integral out of the T-product

→ For the fermionic fields, the statistics modify the T-product:

$$T(\psi(x) \chi(y)) = \theta(x^0 - y^0) \psi(x) \chi(y) - \theta(y^0 - x^0) \chi(y) \psi(x)$$

3.3 Matrix elements of the S-matrix

- Once we have an expression for the evolution operator (the S-matrix), we want to calculate amplitudes of physical processes involving evolution of one state into another (decay, scattering, ...).

DEF | Let $|i\rangle$ and $|f\rangle$ be an initial and final state in the interaction representation. The amplitudes A_{fi} is given by

$$A_{fi} \equiv \langle f | \hat{S} | i \rangle$$

- Since the vacuum and particle states change when interaction is present, we can't construct these states with the $\hat{a}_h$ from the free theory.

↳ We start with something simpler: the vacuum expectation values of T-product of Heisenberg fields, called the Green's functions.

① Two-point Green's function:

- We consider a two-point Green's function because it's the simplest case: $G(x, y) = \langle \Omega | T [\phi_H(x) \phi_H(y)] | \Omega \rangle$ with $|\Omega\rangle$ is the vacuum state of the theory ($\neq$ from $|0\rangle$)

- (See note p. 69) One can show that

$$\langle \Omega | T [\phi_H(x) \phi_H(y)] | \Omega \rangle = \frac{\langle 0 | T [\phi_I(x) \phi_I(y) S] | 0 \rangle}{\langle 0 | S | 0 \rangle}$$

② Generalization:

- The difference between the true vacuum state $|\Omega\rangle$ and the perturbative vacuum $|0\rangle$ can be accounted for by dividing by the matrix element $\langle 0 | S | 0 \rangle$

→ Amplitudes of processes are given by:

$$\begin{aligned}
 \mathcal{A} &= \frac{\langle p_1 \dots p_n | S | k_1 \dots k_m \rangle}{\langle 0 | S | 0 \rangle} \times \left(\text{factors correcting for non-vacuum states} \right) \\
 &= \langle p_1 \dots p_n | S | k_1 \dots k_m \rangle \Big|_{\text{some diagrams removed}}
 \end{aligned}$$

3.4 Calculation of the matrix element. (Wich thm)

→ Since the S-matrix contains the T-ordered products of the field operators, we need to know how to calculate expectation values of such operators:

$$\langle p_1 \dots p_n | T[\varphi(x_1) \dots \varphi(x_r)] | k_1 \dots k_m \rangle$$

⊙ Green's function: vacuum average of T[·]:

DEF The simplest Green's function is called the propagator D_F

$$D_F(x, y) \equiv \langle 0 | T[\varphi(x) \varphi(y)] | 0 \rangle$$

→ Remember the free field is:

$$\varphi(x) = \int \frac{d^3k}{(2\pi)^{3/2} \sqrt{2\omega_k}} \left\{ \hat{a}_k e^{-ikx} + \hat{a}_k^\dagger e^{ikx} \right\}$$

DEF We separate $\varphi(x)$ into the negative-frequency part $\varphi^-(x)$ and its positive-frequency part $\varphi^+(x)$:

$$\varphi^+(x) \equiv \int \frac{d^3k}{(2\pi)^{3/2} \sqrt{2\omega_k}} \hat{a}_k e^{-ikx} \rightarrow \varphi^+ | 0 \rangle = 0$$

$$\varphi^-(x) \equiv \int \frac{d^3k}{(2\pi)^{3/2} \sqrt{2\omega_k}} \hat{a}_k^\dagger e^{ikx} \rightarrow \langle 0 | \varphi^- = 0$$

→ With this notation, let's compute:
 $\rightarrow \varphi(x) \varphi(y) = (\varphi_x^+ + \varphi_x^-) (\varphi_y^+ + \varphi_y^-)$

$$= \varphi_x^+ \varphi_y^+ + \varphi_x^+ \varphi_y^- + \varphi_x^- \varphi_y^+ + \varphi_x^- \varphi_y^-$$

$$= \varphi_x^+ \varphi_y^+ + \varphi_y^- \varphi_x^+ + \varphi_x^- \varphi_y^+ + \varphi_x^- \varphi_y^- + [\varphi_x^+, \varphi_y^-]$$

DEF | When the φ^- ($\hat{a}_k^-$ parts) are before the φ^+ ($\hat{a}_k^+$ parts), it's called a normal order and noted $::$

$$\rightarrow \text{ex: } :\varphi_x \varphi_y: = \varphi_x^+ \varphi_y^+ + \varphi_x^- \varphi_y^- + \varphi_x^- \varphi_y^+ + \varphi_y^- \varphi_x^+$$

$$\hookrightarrow \text{Then we have: } \varphi(x) \varphi(y) = :\varphi(x) \varphi(y): + [\varphi(x)^+, \varphi(y)^-]$$

$$\rightarrow \varphi(y) \varphi(x) = :\varphi(y) \varphi(x): + [\varphi(y)^+, \varphi(x)^-]$$

DEF | Let $\hat{A}, \hat{B}$ be two operators. We define their contraction $\overline{\hat{A} \hat{B}}$ as follow:

$$\overline{\hat{A} \hat{B}} \equiv \hat{A} \hat{B} - :\hat{A} \hat{B}: \quad \overline{\hat{A} \hat{B}} \in \mathbb{C}$$

$\hookrightarrow$ With this notation, we have:

$$\overline{\varphi_x \varphi_y} = [\varphi_x^+, \varphi_y^-] \quad \text{and} \quad \overline{\varphi_y \varphi_x} = [\varphi_y^+, \varphi_x^-] \in \mathbb{C}$$

$\rightarrow$ We can then write:

$$T[\varphi(x) \varphi(y)] = :\varphi(x) \varphi(y): + \overline{\varphi(x) \varphi(y)} \quad \overline{\varphi(x) \varphi(y)} \in \mathbb{C} \text{ number}$$

$\rightarrow$ Coming back to our propagator:

$$\langle 0 | T[\varphi_x \varphi_y] | 0 \rangle = \langle 0 | :\varphi(x) \varphi(y): | 0 \rangle + \overline{\varphi(x) \varphi(y)}$$

$$\Leftrightarrow \mathcal{D}_F(x-y) = \overline{\varphi(x) \varphi(y)} = \begin{cases} [\varphi^+(x), \varphi^-(y)] & \text{at } x^0 > y^0 \\ [\varphi^+(y), \varphi^-(x)] & \text{at } y^0 > x^0 \end{cases}$$

$\hookrightarrow$ This object plays a fundamental role in perturbative calculations.

① Study of the propagator $D_F(x-y)$:

→ $[\varphi^+(x), \varphi^-(y)]$ (case with $x^0 > y^0$)

$$= \int \frac{d^3k d^3q}{(2\pi)^3 2\sqrt{\omega_k \omega_q}} \left\{ \hat{a}_k e^{-i\vec{k}\vec{x}} \hat{a}_q^\dagger e^{i\vec{q}\vec{y}} - \hat{a}_q^\dagger e^{i\vec{q}\vec{y}} \hat{a}_k e^{-i\vec{k}\vec{x}} \right\}$$

$$= \int \frac{d^3k d^3q}{(2\pi)^3 2\sqrt{\omega_k \omega_q}} e^{-i\vec{k}\vec{x} + i\vec{q}\vec{y}} [\hat{a}_k, \hat{a}_q^\dagger] = \delta^3(\vec{k} - \vec{q})$$

$$= \int \frac{d^3k}{(2\pi)^3 2\omega_k} e^{-ik(\vec{x} - \vec{y})}$$

x, y, k are 4 vectors:

$$kx = k_0 x_0 - \vec{k} \cdot \vec{x} = k_0 x_0 - k_i x_i$$

$$= \int \frac{d^3k}{(2\pi)^3 2\omega_k} \exp\{-ik_0(x_0 - y_0) + i\vec{k}(\vec{x} - \vec{y})\}$$

→ $[\varphi^+(y), \varphi^-(x)]$ (case with $y^0 > x^0$)

$$= \int \frac{d^3k}{(2\pi)^3 2\omega_k} \exp\{-ik_0(y_0 - x_0) + i\vec{k}(\vec{y} - \vec{x})\}$$

PROP We can write these 2 expressions as a single one valid all times:

$$D_F(x-y) = \int \frac{d^4k}{(2\pi)^4} \frac{i}{k^2 - m^2 + i\epsilon} e^{-ik(x-y)}$$

Proof:

$$D_F(x-y) = \int \frac{d^3k}{(2\pi)^4} \int_{-\infty}^{\infty} dk_0 \frac{i}{k^2 - m^2 + i\epsilon} e^{-ik^0(x^0 - y^0) + i\vec{k}(\vec{x} - \vec{y})}$$

$$= \int \frac{d^3k}{(2\pi)^4} i e^{i\vec{k}(\vec{x} - \vec{y})} \int_{-\infty}^{\infty} dk_0 \frac{1}{k_0^2 - \omega_k^2 + i\epsilon} e^{-ik^0(x^0 - y^0)}$$

$$\begin{aligned} k^2 - m^2 &= k_0^2 - \vec{k}^2 - m^2 \\ &= k_0^2 - (k^2 + m^2) \\ &= k_0^2 - \omega_k^2 \end{aligned}$$

→ $\frac{1}{k_0^2 - \omega_k^2 + i\epsilon}$ has two poles: $k_0 = \pm \sqrt{\omega_k^2 - i\epsilon} = \pm \sqrt{1 - i\epsilon/\omega_k^2} \cdot \omega_k$

$$\approx \pm \omega_k (1 - i\epsilon/2\omega_k^2)$$

$$\approx \pm \omega_k \mp \frac{i\epsilon}{2\omega_k^2}$$

→ We gonna use the residue theorem to integrate:

$$\oint \gamma \mathcal{L}(z) dz = 2\pi i \sum \text{Res}(\mathcal{L}, a_n)$$

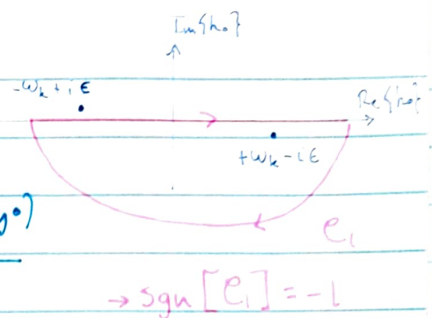
→ Case at $x^0 > y^0$:

↳ $e^{-ik^0(x^0-y^0)} \ll 1$ when: $\text{Im}\{k^0\} \ll -1$

$$\hookrightarrow \int_{-\infty}^{\infty} \frac{dk_0 e^{-ik_0(x^0-y^0)}}{(k_0-\omega_k+i\epsilon)(k_0+\omega_k-i\epsilon)} = (-2\pi i) \frac{e^{-i\omega_k(x^0-y^0)}}{2\omega_k}$$

$$\text{can Res} \left[\frac{e^{-ik_0(x^0-y^0)}}{(k_0-\omega_k+i\epsilon)(k_0+\omega_k-i\epsilon)}, a_k = \omega_k - i\epsilon \right]$$

$$= \lim_{k_0 \rightarrow \omega_k} \left[\frac{e^{-ik_0(x^0-y^0)}}{k_0 + \omega_k - i\epsilon} \right] = \frac{e^{-i\omega_k(x^0-y^0)}}{2\omega_k}$$



$$\begin{aligned} \hookrightarrow \text{We then have } D_F(x-y) &= \int \frac{d^3k}{(2\pi)^4} i e^{ik(\bar{x}-\bar{y})} (-2\pi i) \frac{e^{-i\omega_k(x^0-y^0)}}{2\omega_k} \\ &= \int \frac{d^3k}{(2\pi)^3} \frac{e^{-i\omega_k(x^0-y^0) + ik(\bar{x}-\bar{y})}}{2\omega_k} \end{aligned}$$

→ Case at $y^0 > x^0$:

$$\begin{aligned} \hookrightarrow D_F(x-y) &= \int \frac{d^3k}{(2\pi)^4} i e^{ik(\bar{x}-\bar{y})} \int_{-\infty}^{\infty} \frac{dk_0}{(k_0-\omega_k+i\epsilon)(k_0+\omega_k-i\epsilon)} \\ &= \int \frac{d^3k}{(2\pi)^4} i e^{ik(\bar{x}-\bar{y})} \cdot (2\pi i) \frac{e^{i\omega_k(x^0-y^0)}}{-2\omega_k} \end{aligned}$$

$$= + \int \frac{d^3k}{(2\pi)^3} \frac{-1}{2\omega_k} e^{i\omega_k(x^0-y^0) - ik(\bar{x}-\bar{y})} = \int \frac{d^3k}{(2\pi)^3} \frac{1}{2\omega_k} e^{-i\omega_k(y^0-x^0) + ik(\bar{y}-\bar{x})}$$



⊙ Wick theorem:

→ So far, we calculated $\langle 0 | T[\varphi(x)\varphi(y)] | 0 \rangle = D_F(x-y)$.
How to calculate more complicated ones?

TNM The Wick theorem states:

$$T[\varphi(x_1) \dots \varphi(x_m)] = : \varphi(x_1) \dots \varphi(x_m) : + (\text{all possible contractions}) :$$

$$\rightarrow \text{ex: } T[\hat{\phi}_1 \hat{\phi}_2 \hat{\phi}_3] = : \hat{\phi}_1 \hat{\phi}_2 \hat{\phi}_3 : + : \hat{\phi}_1 : \overline{\hat{\phi}_2 \hat{\phi}_3} + : \hat{\phi}_2 : \overline{\hat{\phi}_1 \hat{\phi}_3} + : \hat{\phi}_3 : \overline{\hat{\phi}_1 \hat{\phi}_2}$$

→ When the vacuum average is taken, only the complete contractions survive (normal order operators vanish when sandwiched with vacuum).

→ ex: $\langle 0 | T[\phi_1 \phi_2 \phi_3 \phi_4] | 0 \rangle = 2 \overbrace{\phi_1 \phi_2 \phi_3 \phi_4} + 2 \overbrace{\phi_1 \phi_2 \phi_4 \phi_3} + 2 \overbrace{\phi_1 \phi_3 \phi_2 \phi_4}$

Proof (of Wick thm) By induction: assume Wick thm for $n-1$ field.
Without loss of generality, assume $x_1^0 \geq x_2^0 \geq \dots \geq x_n^0$

→ $T[\phi_1 \dots \phi_n] = \phi_1 \dots \phi_n = (\phi_1^+ + \phi_1^-) : \{ \phi_2 \dots \phi_n + \text{contractions} \} :$

↳ $\phi_1^+ : \phi_2 \dots \phi_n : = : \phi_2 \dots \phi_n : \phi_1^+ + [\phi_1^+, : \phi_2 \dots \phi_n :]$
 $= : \phi_1^+ \phi_2 \dots \phi_n : + : [\phi_1^+, \phi_2] \phi_3 \dots \phi_n : + : \phi_2 [\phi_1^+, \phi_3] \phi_3 \dots \phi_n :$
 $+ \dots + : \phi_2 \dots \phi_n, [\phi_1^+, \phi_n] :$
 $= : \phi_1^+ \phi_2 \dots \phi_n : + : \overbrace{\phi_1^+ \phi_2 \phi_3 \dots \phi_n} + \dots + : \phi_2 \dots \phi_{n-1} \overbrace{\phi_1^+ \phi_n} :$

→ $T[\phi_1 \dots \phi_n] = : (\phi_1^+ + \phi_1^-) \phi_2 \dots \phi_n + \text{contractions} :$ ▣

⊙ Graphical representation: an example:

→ We consider the $\lambda \phi^4$ theory. We have:

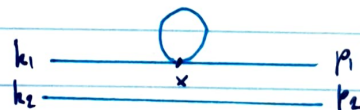
$\langle p_1 p_2 | \hat{S} | k_1 k_2 \rangle = \langle p_1 p_2 | (1 + i\lambda \int d^4x \phi_x^4 + \frac{1}{2} (i\lambda^2) \int d^4x d^4y T(\phi_x^2 \phi_y^2) + \dots$

We'll treat the matrix element by order.

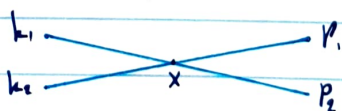
↳ Order 0: $\langle p_1 p_2 | 1 | k_1 k_2 \rangle$ just gives δ -functions

↳ Order 1: $\langle p_1 p_2 | i\lambda \int d^4x \phi_x \phi_x \phi_x \phi_x | k_1 k_2 \rangle$ contracted in all possible ways:

→ $\langle p_1 p_2 | \overbrace{\phi_x \phi_x \phi_x \phi_x} | k_1 k_2 \rangle$



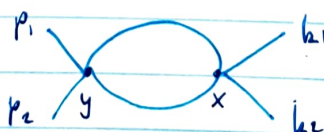
→ $\langle p_1 p_2 | \overbrace{\phi_x \phi_x} \overbrace{\phi_x \phi_x} | k_1 k_2 \rangle$



↳ gives $i\lambda \int d^4x \int e^{-ik_1x} e^{-ik_2x} e^{ip_1x} e^{ip_2x} = i\lambda (2\pi)^4 \delta^4(k_1 + k_2 - p_1 - p_2)$

↳ Order 2: $\frac{1}{2} (i\lambda)^2 \int d^4x d^4y \langle p_1 p_2 | T(\phi_x \phi_x \phi_x \phi_x \phi_y \phi_y \phi_y \phi_y) | k_1 k_2 \rangle$

→ $\langle p_1 p_2 | \overbrace{\phi_x \phi_x \phi_x \phi_x} \overbrace{\phi_y \phi_y \phi_y \phi_y} | k_1 k_2 \rangle$



Rule:

→ Each internal line produces a factor

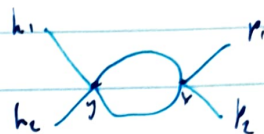
$$\phi(x_1) \phi(x_2) = \int \frac{d^4 q_{12}}{(2\pi)^4} \frac{i}{q_{12}^2 - m^2 + i\epsilon} e^{-i q_{12} (x_1 - x_2)} = D_F(x_1 - x_2)$$

→ Each external line produces a factor

$$\langle k | \phi(x) \Rightarrow e^{ikx} \text{ and } \phi(x) | p \rangle \Rightarrow e^{-ipx}$$

→ For our example, we have:

$$\langle p_1 p_2 | T(\phi_x \phi_x \phi_y \phi_y \phi_z \phi_z \phi_z \phi_z) | k_1 k_2 \rangle \text{ that gives}$$



$$\frac{1}{2!} (i\lambda)^2 \int d^4 x d^4 y \cdot 2 \cdot 4 \cdot 3 \cdot 4 \cdot 3 \cdot 2 e^{-ik_1 x} e^{-ik_2 y} e^{+ip_1 x} e^{+ip_2 y} x$$

$$\cdot \int \frac{d^4 q_1}{(2\pi)^4} \frac{i}{q_1^2 - m^2 + i\epsilon} e^{-i q_1 (x-y)} \cdot \int \frac{d^4 q_2}{(2\pi)^4} \frac{i}{q_2^2 - m^2 + i\epsilon} e^{-i q_2 (x-y)}$$

$$= \frac{(4!)^2}{2!} (i\lambda)^2 \int \frac{d^4 q_1}{(2\pi)^4} \frac{d^4 q_2}{(2\pi)^4} \left\{ (2\pi)^4 \delta^4(p_2 + p_1 - q_1 - q_2) \cdot (2\pi)^4 \delta^4(q_1 + q_2 - k_1 - k_2) \right.$$

$$\times \frac{i}{q_1^2 - m^2 + i\epsilon} \cdot \frac{i}{q_2^2 - m^2 + i\epsilon} \Big\}$$

$$= \frac{(4!)^2}{2} (i\lambda)^2 (2\pi)^4 \delta^4(p_2 + p_1 - k_1 - k_2) \int \frac{d^4 q_1}{(2\pi)^4} \frac{i}{q_1^2 - m^2 + i\epsilon} \frac{i}{(k_1 + k_2 - q_1)^2 - m^2 + i\epsilon}$$

↳ It diverges!

Rules

(Feynman)

①: Draw all possible diagrams

②: Vertex $\rightarrow (-i\lambda) (2\pi)^4 \delta^4(\sum k_i)$ ③: Each internal line: $\int d^4 k / (2\pi)^4 \cdot i / (k^2 - m^2 + i\epsilon)$

④: Each external line: 1

3.5 Desintegration of a massive particle

○ From the S-matrix to probabilities:

→ Normalization of states $|k\rangle$: how many particles does one $|k\rangle$ contain?

$$\begin{aligned}\langle k|N|k\rangle &= \langle k|k\rangle \\ &= \langle 0| (2\pi)^{3/2} \sqrt{2\omega_k} \hat{a}_k \cdot (2\pi)^{1/2} \sqrt{2\omega_k} \hat{a}_k^\dagger |0\rangle \\ &= (2\pi)^3 2\omega_k \delta^3(k-k) = (2\pi)^3 2\omega_k \delta^3(0)\end{aligned}$$

↳ Meaning of $\delta^3(0)$?

$$\rightarrow \delta^3(k) = \int \frac{d^3x}{(2\pi)^3} e^{ikx}$$

$$\rightarrow \delta^3(0) = \int \frac{d^3x}{(2\pi)^3} = \frac{V}{(2\pi)^3}$$

↳ We get $\langle k|k\rangle = 2\omega_k V$, it means that our states contain $2\omega_k V$ particles.

→ The state containing 1 particle is

$$|\phi_k\rangle \equiv \frac{1}{\sqrt{2\omega_k} \sqrt{V}} |k\rangle$$

↳ We can create a state with 2 particles:

$$|\phi_{p_1} \phi_{p_2}\rangle = \frac{1}{\sqrt{2\omega_{p_1} V}} \frac{1}{\sqrt{2\omega_{p_2} V}} |p_1 p_2\rangle$$

→ From QM, the amplitude of $\phi_k \rightarrow \phi_{p_1} + \phi_{p_2}$ is given by:

$$\langle \phi_{p_1} \phi_{p_2} | S | \phi_k \rangle$$

and the probability is $|\langle \phi_{p_1} \phi_{p_2} | S | \phi_k \rangle|^2$

↳ But we have a problem:

$$|\langle \phi_{p_1} \phi_{p_2} | S | \phi_k \rangle|^2 = |\langle p_1 p_2 | S | k \rangle|^2 \frac{(2\omega_{p_1})^{-1}}{V} \frac{(2\omega_{p_2})^{-1}}{V} \frac{(2\omega_k)^{-1}}{V} \xrightarrow{V \rightarrow \infty} 0$$

→ We need to look at the probability to decay in a state with momenta of particles in the intervals $d^3\vec{p}_1$ and $d^3\vec{p}_2$

From QM, the # of state in d^3p is:

$$\frac{d^3p \cdot V}{(2\pi)^3}$$

↳ The probability of decay in a state with momenta $[\vec{p}_1; \vec{p}_1 + \Delta \vec{p}_1]$, $[\vec{p}_2; \vec{p}_2 + \Delta \vec{p}_2]$ is

$$dP = |\langle \varphi_{p_1} \varphi_{p_2} | S | \varphi_k \rangle|^2 \cdot \frac{d\vec{p}_1 V}{(2\pi)^3} \cdot \frac{d\vec{p}_2 V}{(2\pi)^3}$$

$$= \frac{1}{2\omega_k V} \cdot \frac{d^3 p_1}{(2\pi)^3 2\omega_{p_1}} \cdot \frac{d^3 p_2}{(2\pi)^3 2\omega_{p_2}} |\langle p_1 p_2 | S | k \rangle|^2$$

→ Since we know that $|\langle p_1 p_2 | S | k \rangle|^2$ always contain the factor $(2\pi)^4 \delta^4(\sum k - \sum p)$ to ensure energy-momentum conservation, we can write:

$$\langle p_1 \dots p_n | \hat{S} | k_1 \dots k_n \rangle \equiv i (2\pi)^4 \delta^4(\sum k - \sum p) \cdot \mathcal{M}(p_1 \dots p_n)$$

↳ We have then:

$$|\langle p_1 p_2 | \hat{S} | k \rangle|^2 = (2\pi)^8 \delta^4(k - p_1 - p_2) \delta^4(k - p_1 - p_2) |\mathcal{M}|^2$$

$$= (2\pi)^8 \delta^4(k - p_1 - p_2) \delta^4(0) |\mathcal{M}|^2$$

$$= (2\pi)^8 \delta^4(k - p_1 - p_2) \frac{V \cdot T}{(2\pi)^4} |\mathcal{M}|^2$$

$$= (2\pi)^4 V \cdot T \cdot \delta^4(k - p_1 - p_2) |\mathcal{M}|^2$$

$$\delta^4(0) = \frac{V}{(2\pi)^3}$$

→ Finally, we have:

prop |
$$d\Gamma \equiv \frac{dP}{T} = (2\pi)^4 \delta^4(k - \sum p) \frac{1}{2\omega_k} |\mathcal{M}|^2 \cdot \frac{d^3 p_1}{(2\pi)^3 2\omega_{p_1}} \cdot \frac{d^3 p_2}{(2\pi)^3 2\omega_{p_2}}$$

where Γ is the decay rate

⊙ Interpretation of Γ :

→ Γ is the probability of decay in a unit time.

→ If you take a large number of particles N , then their number depends on time according to:

$$N(t) = N_0 e^{-\Gamma t} = N_0 e^{-t/\tau}$$

↳ We see that

$$\Gamma = \frac{1}{\tau}$$

② Total decay rate:

→ Setting $\omega_k = M$, we have:

$$\Gamma = \int \frac{dP}{T} = \frac{1}{8\pi^2 M} \int \frac{d^3 p_1 d^3 p_2}{2\omega_{p_1} 2\omega_{p_2}} |CM|^2 \delta^4(k - p_1 - p_2)$$

Using $\delta^4(k - p_1 - p_2) = \delta(M - \omega_{p_1} - \omega_{p_2}) \cdot \delta^3(\vec{p}_1 + \vec{p}_2)$, we integrate:

$$\Gamma = \frac{1}{32\pi^2 M} \int \frac{d^3 p_1}{\omega_1 \omega_2} |CM|^2 \delta(M - \omega_1 - \omega_2)$$

$$\text{with } \omega_i = \sqrt{m_i^2 + p_i^2}$$

→ We use $\int dx \delta(x) \delta[g(x)] = \sum_{\text{root}} \frac{\delta(x_i)}{|g'(x_i)|}$

$$\text{Here, } g(p_i) = M - \sqrt{m_1^2 + p_1^2} - \sqrt{m_2^2 + p_1^2}$$

$$\hookrightarrow \partial_{p_i} g = -\frac{p_1}{\omega_1} - \frac{p_1}{\omega_2} = -p_1 \frac{(\omega_1 + \omega_2)}{\omega_1 \omega_2}$$

$$\frac{1}{|g'|} = \frac{\omega_1 \omega_2}{p_1 (\omega_1 + \omega_2)} = \frac{\omega_1 \omega_2}{p_1 M}$$

Where did the $\delta(x)$ go?
+ why $M_{\text{new}} = \langle E_{\text{initial}} \rangle$

→ Going in spherical coord., we have: $d^3 p = p^2 dp d\Omega$

$$d\Gamma = \frac{1}{32\pi^2 M} |CM|^2 \frac{p^2 dp d\Omega}{\omega_1 \omega_2} \cdot \frac{\omega_1 \omega_2}{p_1 M}$$

$$d\Gamma = \frac{p_1 d\Omega}{32\pi^2 M^2} |CM|^2$$

→ Case where $m_1 = m_2$, we have:

$$M = 2\omega = 2\sqrt{m^2 + p^2} \Leftrightarrow M^2 = 4(m^2 + p^2) \Rightarrow p_1 = \sqrt{\frac{M^2}{4} - m^2}$$

$$\hookrightarrow d\Gamma = \frac{1}{32\pi^2 M^2} \cdot \sqrt{\frac{M^2}{4} - m^2} d\Omega |CM|^2$$

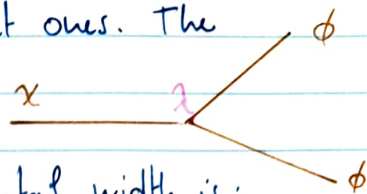
$$= \frac{1}{64\pi^2 M} \sqrt{1 - \frac{4m^2}{M^2}} |CM|^2 d\Omega$$

→ If $|CM|$ is not dependent on the angle, $d\Omega \rightarrow 4\pi$. We get

$$\Gamma = \frac{1}{16\pi M} \sqrt{1 - \frac{4m^2}{M^2}} |CM|^2$$

② Particular examples:

→ Simplest case: heavy scalar into 2 light ones. The interaction term in this case is $\lambda \chi \phi^2$



↳ We have $|M|^2 = \lambda^2$. Therefore, the total width is:

$$\Gamma = \frac{\lambda^2}{32\pi M_\chi} \sqrt{1 - \frac{4m_\phi^2}{M_\chi^2}} \approx \frac{\lambda^2}{32\pi M_\chi}$$

3.6 Scattering and Cross section

DEF For a single target, the cross section σ is the area in the plane perpendicular to the incoming flux from which the particles are scattered.

→ For a beam of particles with density n_A and velocity v_A , the # of event in a time interval T is

$$N_{ev} = n_A \cdot v_A \cdot \sigma T$$

$$\Leftrightarrow \sigma = \frac{N_{ev}}{n_A v_A \cdot T}$$

→ Let's consider a normalized initial state $|\varphi_A \varphi_B\rangle$

$$|\varphi_A \varphi_B\rangle = \frac{1}{\sqrt{2\omega_A V}} \frac{1}{\sqrt{2\omega_B V}} |k_A k_B\rangle = \frac{(2\pi)^{3/2}}{\sqrt{V}} \frac{(2\pi)^{3/2}}{\sqrt{V}} a_{k_A}^\dagger a_{k_B}^\dagger |0\rangle$$

$$\hookrightarrow N_{ev} = |\langle \ell | \hat{S} | \varphi_A \varphi_B \rangle|^2 = \frac{|\langle \ell | \hat{S} | k_A k_B \rangle|^2}{2\omega_A V \cdot 2\omega_B V}$$

↳ Since it's normalized (1 particle per volume V), $n_A = 1/V$

$$\hookrightarrow \text{We get: } \sigma = \frac{|\langle \ell | \hat{S} | k_A k_B \rangle|^2}{4\omega_A \omega_B V^2} \cdot \frac{V}{v_A \cdot T} = \frac{|\langle \ell | \hat{S} | k_A k_B \rangle|^2}{4\omega_A \omega_B v_A T} \cdot \frac{1}{V} \quad \text{Not well defined.}$$

→ We need to consider the probability of the transition into a state with momenta between $\vec{p}_i$ and $\vec{p}_i + d\vec{p}_i$. Then

$$d\sigma = \frac{|\langle p_1 \dots p_n | \hat{S} | k_A k_B \rangle|^2}{2\omega_1 \dots 2\omega_n \cdot 2\omega_A \cdot 2\omega_B} \cdot \frac{1}{v_A V T} \cdot \frac{d\vec{p}_1}{(2\pi)^3} \dots \frac{d\vec{p}_n}{(2\pi)^3}$$

↳ Since our matrix element always contain a factor δ^4 , we can write

$$\langle p_1 \dots p_n | \hat{S} | k_A k_B \rangle = i (2\pi)^4 \delta^4(\Sigma k - \Sigma p) \mathcal{M}(p_1 \dots k_B)$$

choice

$$\text{Then } |\langle p_1 \dots p_n | \hat{S} | k_A k_B \rangle|^2 = (2\pi)^8 \delta^4(\Sigma k - \Sigma p) \delta^4(\Sigma k - \Sigma p) |\mathcal{M}|^2 \\ = (2\pi)^8 \delta^4(\Sigma k - \Sigma p) V.T |\mathcal{M}|^2$$

$$\delta^4(0) = \frac{VT}{(2\pi)^4}$$

$$\rightarrow \text{We find: } d\sigma = \frac{|\mathcal{M}|^2 (2\pi)^4 \delta^4(\Sigma p - \Sigma k)}{4\omega_A \omega_B t_A} \frac{d^3 p_1 \dots d^3 p_n}{(2\pi)^3 2\omega_1 (2\pi)^3 2\omega_n}$$

→ In the frame where B is at rest:

$$k_A = (\omega_A, \vec{k}_A); k_B = (m_B, 0); k_A k_B = \omega_A m_B = \omega_A \omega_B$$

$$\text{Then, } (k_A k_B)^2 - m_A^2 m_B^2 = \omega_A^2 \omega_B^2 - m_A^2 m_B^2$$

▽

$$= m_B^2 (\omega_A^2 - m_A^2) = m_B^2 m_A^2 \left(\frac{1}{1 - v_A^2} - 1 \right)$$

$$= m_B^2 m_A^2 \frac{v_A^2}{1 - v_A^2} = m_B^2 E_A^2 v_A^2$$

We have

$$d\sigma = \frac{|\mathcal{M}|^2}{4 \sqrt{(k_A k_B)^2 - m_A^2 m_B^2}} (2\pi)^4 \delta^4(\Sigma k - \Sigma p) \frac{d^3 p_1 \dots d^3 p_n}{(2\pi)^3 2\omega_1 (2\pi)^3 2\omega_n}$$